

### Solution to Exercise 2.1.

We have

$$\begin{aligned}m_{\text{app}} &= -2.5 \log I + K_1 \\ &= -2.5 \log L + 2.5 \log (4\pi d^2) + K_1 \\ &= M_{\text{abs}} - K_2 + 2.5 \log (4\pi) + 5 \log d + K_1,\end{aligned}$$

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which gives the required relation (2.8) when we choose

$$K_2 = K_1 + 2.5 \log (4\pi) + 5.$$

### Solution to Exercise 2.2.

We have

$$\begin{aligned}I_B &= \frac{L_B}{4\pi d^2}, & I_U &= \frac{L_U}{4\pi d^2}, \\ & & I_B &= \frac{L_B}{4\pi d^2}, & I_U &= \frac{L_U}{4\pi d^2}, \\ I_B &= \frac{L_B}{4\pi d^2}, & I_U &= \frac{L_U}{4\pi d^2},\end{aligned}$$

$$\frac{I_B}{I_U} = \frac{L_B}{L_U},$$

$$\frac{I_B}{I_U} = \frac{L_B}{L_U} \frac{I_B}{I_U} = \frac{L_B}{L_U},$$

And hence  $U - B$  is independent on distance  $d$ ; the same with other colour indices.

### Solution to Exercise 2.3.

We introduce

$$x = \frac{hc}{\lambda kT}$$

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as new independent variable, which gives

$$\begin{aligned}
 I_{\text{bol}} &= \int_0^{\infty} I_{\lambda} d\lambda \\
 &= 2\pi hc^2 \left(\frac{kT}{hc}\right)^4 \left(\frac{R}{d}\right)^2 \int_0^{\infty} \frac{x^3}{e^x - 1} dx \\
 &= \sigma T^4 \left(\frac{R}{d}\right)^2
 \end{aligned}$$

$$\begin{aligned}
 I_{\text{bol}} &= \int_0^{\infty} I_{\lambda} d\lambda & I_{\text{bol}} &= \int_0^{\infty} I_{\lambda} d\lambda \\
 &= 2\pi hc^2 \left(\frac{kT}{hc}\right)^4 \left(\frac{R}{d}\right)^2 \int_0^{\infty} \frac{x^3}{e^x - 1} dx & &= 2\pi hc^2 \left(\frac{kT}{hc}\right)^4 \left(\frac{R}{d}\right)^2 \int_0^{\infty} \frac{x^3}{e^x - 1} dx \\
 &= \sigma T^4 \left(\frac{R}{d}\right)^2 & &= \sigma T^4 \left(\frac{R}{d}\right)^2
 \end{aligned}$$

where

$$\sigma = 2\pi hc^2 \left(\frac{k}{hc}\right)^4 \int_0^{\infty} \frac{x^3}{e^x - 1} dx.$$

$$\sigma = 2\pi hc^2 \left(\frac{k}{hc}\right)^4 \int_0^{\infty} \frac{x^3}{e^x - 1} dx \quad \sigma = 2\pi hc^2 \left(\frac{k}{hc}\right)^4 \int_0^{\infty} \frac{x^3}{e^x - 1} dx.$$

**Solution to Exercise 2.4.**

Parallax

$$p'' = \frac{1 \text{ pc}}{d(\text{pc})} = \frac{3.1 \times 10^{16} \text{ m}}{1.2 \times 10^{17} \text{ m}} = 0.258$$

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Distance

$$d(\text{pc}) = \frac{1 \text{ pc}}{p''} = 2 \text{ pc} = 6.2 \times 10^{16} \text{ m}$$

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True ratio of brightness is

$$10 \times \left( \frac{6.2}{12} \right)^2 = 2.67 .$$

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**Solution to Exercise 2.5.**

a)

$$\begin{aligned} M_{\text{abs}} &= m_{\text{app}} - 5 \log d(\text{pc}) + 5 \\ &= 7.62 - 10 + 5 = 2.62 \end{aligned}$$

$$\begin{aligned} M_{\text{abs}} &= m_{\text{app}} - 5 \log d(\text{pc}) + 5 & M_{\text{abs}} &= m_{\text{app}} - 5 \log d(\text{pc}) + 5 \\ &= 7.62 - 10 + 5 = 2.62 & &= 7.62 - 10 + 5 = 2.62 \end{aligned}$$

b) We have an equation

$$M_{\text{abs}} - M_{\text{abs},\odot} = -2.5 \log \frac{L}{L_{\odot}} = 2.62 - 4.62 = -2.00$$

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from which

$$\log \frac{L}{L_{\odot}} = 0.8, \quad \frac{L}{L_{\odot}} = 10^{0.8} = 6.3,$$

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$$L = 6.3 \times 3.86 \times 10^{26} \text{ J/s} = 2.43 \times 10^{27} \text{ J/s}$$

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c)

$$R^2 = \frac{L}{4\pi\sigma T_{\text{eff}}^4} = \frac{2.43 \times 10^{27}}{4\pi \times 5.67 \times 10^{-8} \times 5780^4}$$
$$= 3.06 \times 10^{18} (\text{m}^2), \quad R = 1.75 \times 10^9 \text{ m}$$

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$$= 3.06 \times 10^{18} (\text{m}^2), \quad R = 1.75 \times 10^9 \text{m}$$

Second star:

$$L_2 = 100^2 L_1,$$

$$M_{\text{abs},2} = M_{\text{abs},1} - 2.5 \log \frac{L_2}{L_1} = 2.62 - 10 = -7.38,$$

$$m_{\text{app},2} = M_{\text{abs},2} + 5 \log d_2 - 5 = 13.12,$$

$$5 \log d_2 = 13.12 + 5 + 7.38 = 25.5,$$

$$\log d_2 = 5.1, \quad d_2 = 10^{5.1} = 1.25 \times 10^5 (\text{pc}),$$

$$p = 8.0 \times 10^{-6} \text{arc sec}$$

**Solution to Exercise 2.6.**

$$m_{\text{app}} = -2.5 \log (\text{count}) + \text{const}$$

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In R filter:

$$m_{1R} - m_{2R} = -2.5 \log \frac{23456}{58919} = 1.000,$$

$$m_{2R} = m_{1R} - 1.0 = 12.1 - 1.0 = 11.1$$

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$$m_{2R} = m_{1R} - 1.0 = 12.1 - 1.0 = 11.1$$

In B filter:

$$m_{1B} - m_{2B} = -2.5 \log \frac{20954}{49405} = 0.932$$

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For solar-type star,

$$B_1 - R_1 = 1.17$$

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(same as for the Sun), hence

$$B_1 \equiv m_{1B} = 1.17 + 12.1 = 13.27$$

$$B_1 \equiv m_{1B} = 1.17 + 12.1 = 13.27 \quad B_1 \equiv m_{1B} = 1.17 + 12.1 = 13.27$$

For other star,

$$B_2 \equiv m_{2B} = m_{1B} - 0.932 = 12.34,$$

$$B_2 - R_2 \equiv m_{2B} - m_{2R} = 12.34 - 11.1 = 1.24$$

$$B_2 \equiv m_{2B} = m_{1B} - 0.932 = 12.34,$$

$$B_2 - R_2 \equiv m_{2B} - m_{2R} = 12.34 - 11.1 = 1.24$$

$$B_2 \equiv m_{2B} = m_{1B} - 0.932 = 12.34,$$

$$B_2 - R_2 \equiv m_{2B} - m_{2R} = 12.34 - 11.1 = 1.24$$

For the second star, **B - R** is bigger, hence it is redder.

### Solution to Exercise 2.7.

a)

$$M_{\text{abs}} = 4.62 - 2.5 \log \frac{L}{L_{\odot}}$$

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Reigel:

$$-6.77 = 4.62 - 2.5 \log \frac{L_1}{L_{\odot}},$$

$$\log \frac{L_1}{L_{\odot}} = \frac{11.39}{2.5} = 4.56, \quad \frac{L_1}{L_{\odot}} = 3.6 \times 10^4$$



$$-6.77 = 4.62 - 2.5 \log \frac{L_1}{L_{\odot}},$$

$$\log \frac{L_1}{L_{\odot}} = \frac{11.39}{2.5} = 4.56, \quad \frac{L_1}{L_{\odot}} = 3.6 \times 10^4$$

$$-6.77 = 4.62 - 2.5 \log \frac{L_1}{L_{\odot}},$$

$$\log \frac{L_1}{L_{\odot}} = \frac{11.39}{2.5} = 4.56, \quad \frac{L_1}{L_{\odot}} = 3.6 \times 10^4$$

$\beta$  Canis Majoris:

$$-1.33 = 4.62 - 2.5 \log \frac{L_2}{L_{\odot}},$$

$$\log \frac{L_2}{L_{\odot}} = 2.38, \quad \frac{L_2}{L_{\odot}} = 240$$

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$$-1.33 = 4.62 - 2.5 \log \frac{L_2}{L_{\odot}},$$

$$\log \frac{L_2}{L_{\odot}} = 2.38, \quad \frac{L_2}{L_{\odot}} = 240$$

b)

$$\frac{L_1}{L_2} = \left( \frac{R_1}{R_2} \right)^2,$$

$$\frac{3.6 \times 10^4}{240} = \left( \frac{R_1}{R_2} \right)^2, \quad \frac{R_1}{R_2} = 12.$$

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$$\frac{3.6 \times 10^4}{240} = \left( \frac{R_1}{R_2} \right)^2, \quad \frac{R_1}{R_2} = 12, \quad \frac{3.6 \times 10^4}{240} = \left( \frac{R_1}{R_2} \right)^2, \quad \frac{R_1}{R_2} = 12.$$

**Solution to Exercise 2.8.**

We have

$$L \propto T^6$$

$$L \propto T^6, \quad L \propto T^6$$

And also

$$L \propto R^2 T^4,$$

$$L \propto R^2 T^4, \quad L \propto R^2 T^4,$$

hence

$$T^6 \propto R^2 T^4, \quad T^2 \propto R^2, \quad T \propto R.$$